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Entanglement Entropy and AdS/CFT Correspondence

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1 Introduction

In quantum mechanical systems, it is very useful to measure how entangled ground states are. An important measure is the entanglement entropy.

[Various Applications]

- Quantum Information and quantum computing
- Condensed matter physics (EE = <u>a new order parameter?</u>)
- Black hole physics (EE = the <u>origin of BH entropy</u>?)

Why not, in String theory?

Definition of entanglement entropy

Divide a given quantum system into two parts A and B. Then the total Hilbert space becomes factorized

$$H_{tot} = H_A \otimes H_B \quad .$$

We define the reduced density matrix ρ_A for A by

$$\rho_A = \mathrm{Tr}_B \rho_{tot} ,$$

taking trace over the Hilbert space of B . Now the entanglement entropy S_A is defined by the von Neumann entropy

$$S_A = -\mathrm{Tr}_A \ \rho_A \ \mathrm{log}\rho_A$$

In this talk we consider the entanglement entropy in quantum field theories on (d+1) dim. spacetime

 $M = R_t \times N.$

Then, we divide N into A and B by specifying the boundary $\partial A = \partial B \subset N$.



E.E. is the entropy for an observer who is only accessible to the subsystem A and not to B.

This setup is very similar to black holes !This property suggests its gravity dual description.

Motivated by this, we would like to apply AdS/CFT to find a holographic dual of entanglement entropy.

Other Useful Properties

- (i) E.E. is a sort of a `non-local version of correlation functions'. (cf. Wilson loops)
- (ii) E.E. is proportional to the degrees of freedom.
 It is non-vanishing even at zero temperature.
 (~ a generalization of the central charge)

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② Holographic Entanglement Entropy Formula

Consider the AdS/CFT in Poincare Coordinate.



Our Holographic Formula

- (1) Divide the space N is into A and B.
- (2) Extend their boundary ∂A to the entire AdS space. This defines a d dimensional surface.
- (3) Pick up a minimal area surface and call this γ_A .
- (4) The E.E. is given by naively applying the Bekenstein-Hawking formula

$$S_A = \frac{\operatorname{Area}(\gamma_A)}{4G_N^{(d+2)}}.$$

as if γ_A were an event horizon.



Motivation of this formula

Here we employ the global coordinate of AdS space and take its time slice at $t=t_0$.



Leading divergence

For a generic choice of $\,\gamma_{\rm A}\,$, $\,$ a basic property of AdS gives

Area
$$(\gamma_{\rm A}) \sim R^d \cdot \frac{\operatorname{Area}(\partial \gamma_{\rm A})}{a^{d-1}}$$
 + (subleading terms),

where R is the AdS radius.

Because $\partial \gamma_A = \partial A$, we find

$$S_A \sim \frac{\operatorname{Area}(\partial A)}{a^{d-1}}$$
 + (subleading terms).

This agrees with the 'area law' of E.E. in QFTs.

Bombelli-Koul-Lee-Sorkin 86', Srednicki 93'

A proof of the holographic formula via GKP-Witten relation [Fursaev hep-th/0606184]

In the CFT side, the (negative) deficit angle $2\pi(1-n)$ is localized on ∂A .



Naturally, it can be extended inside the bulk AdS by solving Einstein equation. We call γ_A this extended surface.

Let us apply the GKP-Witten formula $Z_{CFT} = e^{-S_{Gravity}(\phi_i)}$ in this background with the deficit angle $\Delta \varphi = 2\pi(1-n)$. The curvature is delta functionally localized on the deficit angle surface:



3 Several Examples

(3-1) AdS3/CFT2 Case

For any 2D CFTs, we can exactly compute E.E. for arbitrary choices of the subsystem A. [J.Cardy 2004']

In the dual gravity side, the minimal surfaces are now equivalent to the space-like geodesics.

Thus we can compare both results directly. We checked that they completely agree with each other.

$\frac{\text{Consider the setup}}{\text{Consider the setup}} \left\{ \begin{array}{l} \text{total system = An infinite line} \\ \text{subsystem A = An interval of length } l \end{array} \right.$

The holographic formula leads to $S_{A} = \frac{|\gamma_{A}|}{4G_{M}^{(3)}} = \frac{c}{3}\log\left(\frac{l}{a}\right), \qquad \substack{l \neq \gamma_{A} \\ z = a} \xrightarrow{\gamma_{A}} z$

where we employed the celebrated expression $c = \frac{3R}{2G_N^{(3)}}$. of the central charge.

This result from AdS3 perfectly agrees with the known formula in 2D CFT. (remember a is the UV cutoff)

<u>At Finite temperature $T=1/\beta$ </u>

In this case, the dual gravity background is the BTZ black Hole and the geodesic distance can be found exactly.

This again reproduces the known formula at finite temperature.

$$S_A = \frac{c}{3} \log \left(\frac{\beta}{a} \sinh \left(\frac{\pi l}{\beta} \right) \right)$$

At higher temperature, it includes the thermal contribution

$$S_A \approx \frac{\pi}{3} c \, lT$$

Geometric Interpretation(i) Small A(ii) Large A



When A is large (i.e. high temperature), γ_A wraps a part of horizon. This leads to the thermal contribution $S_A \approx (\pi/3)c \, lT$ to the entanglement entropy.

(3-2) Higher Dimensional Case

Now we compute the holographic E.E. in the Poincare metric dual to a CFT on $R^{1,d}$. To obtain analytical results, we assume that the subsystem A = an infinite strip.



Entanglement Entropy in N=4 Super Yang-Mills

Consider the setup of type IIB string on $AdS_5 \times S^5$, which is dual to 4D N=4 SU(N) super Yang-Mills theory.

In this case, the holographic formula leads to (this corresponds to the strongly coupled Yang-Mills)



Now, we would like to compare this with <u>free</u> Yang-Mills result.

The AdS results numerically reads

$$S_A^{AdS} = K' \cdot \frac{N^2 L^2}{a^2} - 0.051 \cdot \frac{N^2 L^2}{l^2}.$$

On the other hand, the free Yang-Mills result becomes

$$S_{A}^{freeCFT} = K \cdot \frac{N^{2}L^{2}}{a^{2}} - 0.078 \cdot \frac{N^{2}L^{2}}{l^{2}}$$

The order one deviation is expected since the AdS result corresponds to the strongly coupled Yang-Mills. Thus this is regarded as a semi-quantitative matching.

[cf. 4/3 problem in thermal entropy, Gubser-Klebanov-Peet 96']

(3-3) Black hole entropy and Entanglement Entropy [Emparan hep-th/0603081]

<u>Claim</u>

Entropy of 3D quantum black hole = Entanglement Entropy



(3-4) Non-supersymmetric AdS bubble backgrounds

As a final example we consider the asymptotic AdS backgrounds which are obtained from the double Wick rotation of R-charged AdS black holes.

They are a series of static bubble solutions dual to the N=4 SYM with SUSY breaking twisted boundary conditions.

At a specific value of the twist parameter, the N=4 SUSY is recovered.

→ We can break the SUSY very softly!

The entanglement entropies computed in the free Yang-Mills and the AdS gravity agree nicely!



This is a quantitative evidence for AdS/CFT in non-SUSY b.g.!

(4) Conclusions

• We have given a holographic interpretation of entanglement entropy via AdS/CFT duality.

Generalized Area Law

Minimal Surface Area in AdS = Entanglement Entropy in CFT

This strongly suggests

Minimal Surface γ_A in AdS = Holographic Screen for A

It clarifies the issue of *the locality in AdS/CFT*:

Which place in AdS encodes the information of the dual CFT restricted to a given region??



AdS3 in Poincare Coodinate

Happy Birthday to Yoneya-san !

